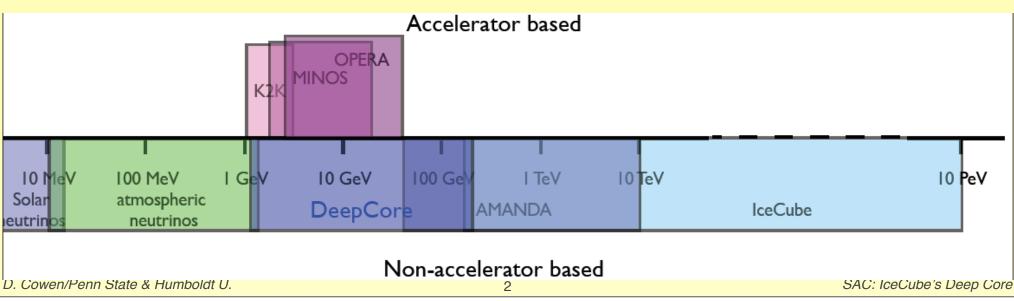
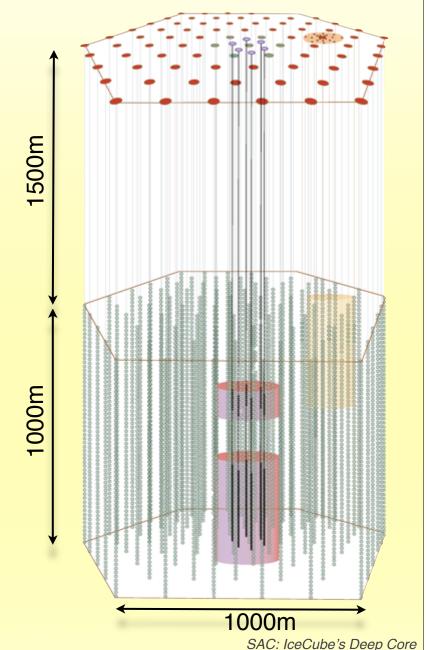
#### Low $E_v \& WIMPs$ with DeepCore

- Deep Core design
- Initial studies of first DeepCore string
- Predicted DeepCore performance
- Status of low energy analyses
  - WIMPs
  - Neutrino oscillations
- Enhancing DeepCore with two additional strings ("79&80")

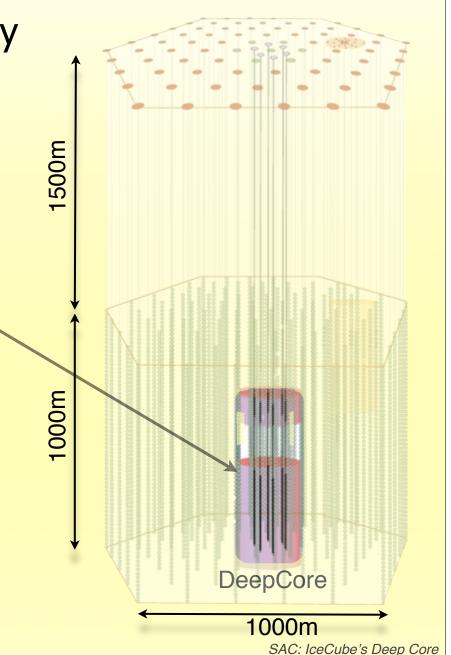
- IceCube: Optimized for 1 TeV <  $E_v$  < 1 PeV
  - "map v sky": search for high energy astrophysical v's from GRBs, SNe, AGN; search for WIMP dark matter,...
  - also use these v's to study fundamental v properties: oscillations, decay,...
- DeepCore: Sensitivity as low as  $E_v \sim 10 \text{ GeV}$ 
  - also "map low-E<sub>v</sub> sky"; search for WIMP dark matter
    - access more potential sources; lower mass WIMPs
  - also study fundamental v properties
    - but with a <u>guaranteed</u> source: atmospheric v's



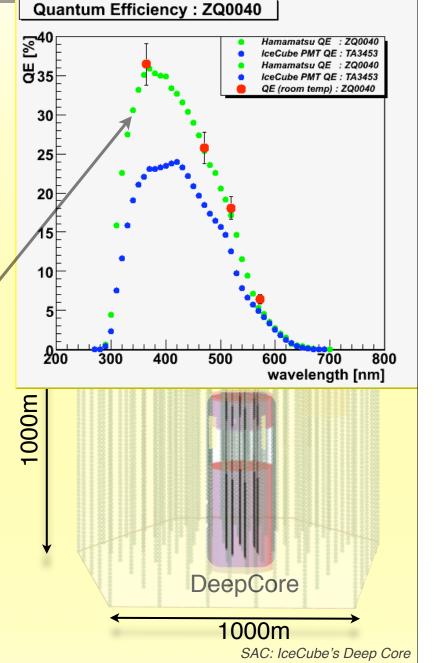
- DeepCore extends low energy reach by ~1 order of mag.
- Design:
  - 6 special strings + 7 nearby IceCube strings
    - 72m interstring spacing
    - 7m vertical spacing
      - 10x higher DOM density
  - high QE Hamamatsu PMTs
  - primarily in clearest deep ice
    - $\lambda_{eff} \sim 40-50m$
  - surrounding strings serve as highly effective active veto
- Funding for hardware (PMTs, strings, etc.) from Europe



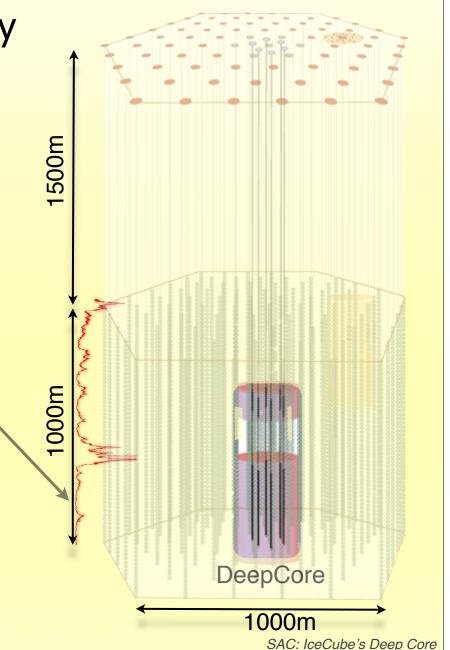
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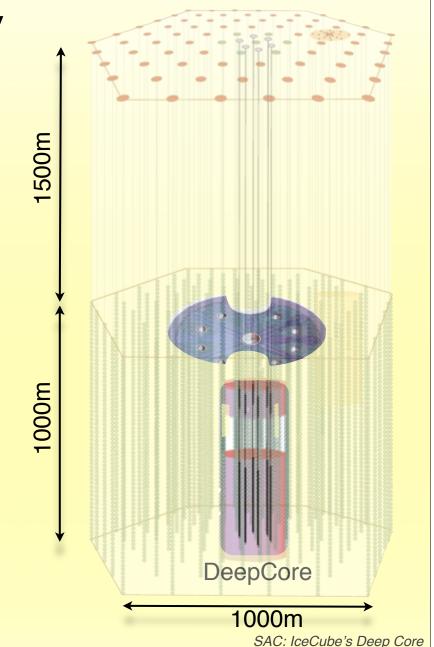
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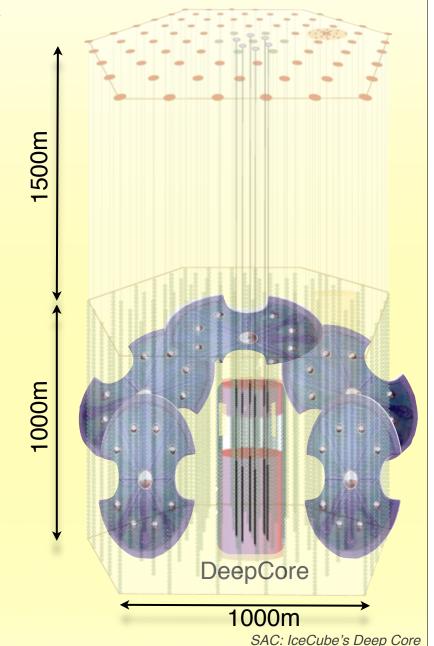
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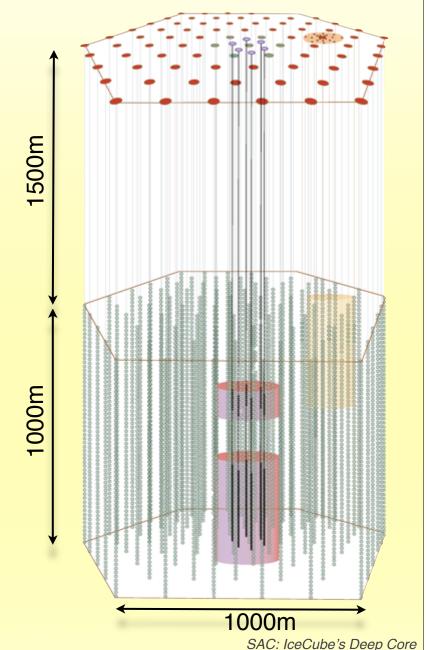
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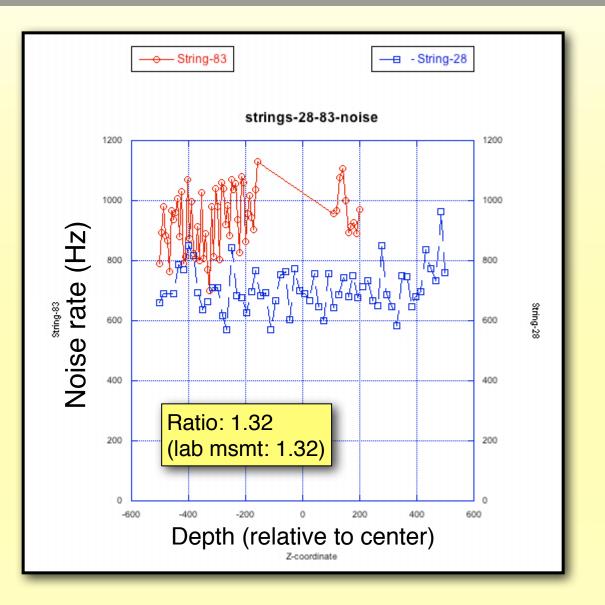


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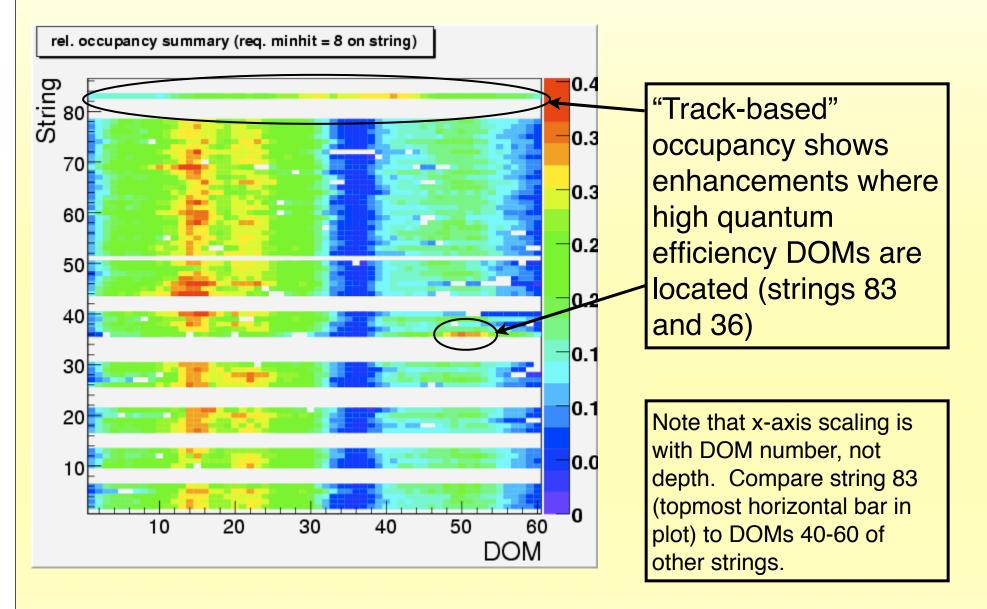


# DeepCore Performance

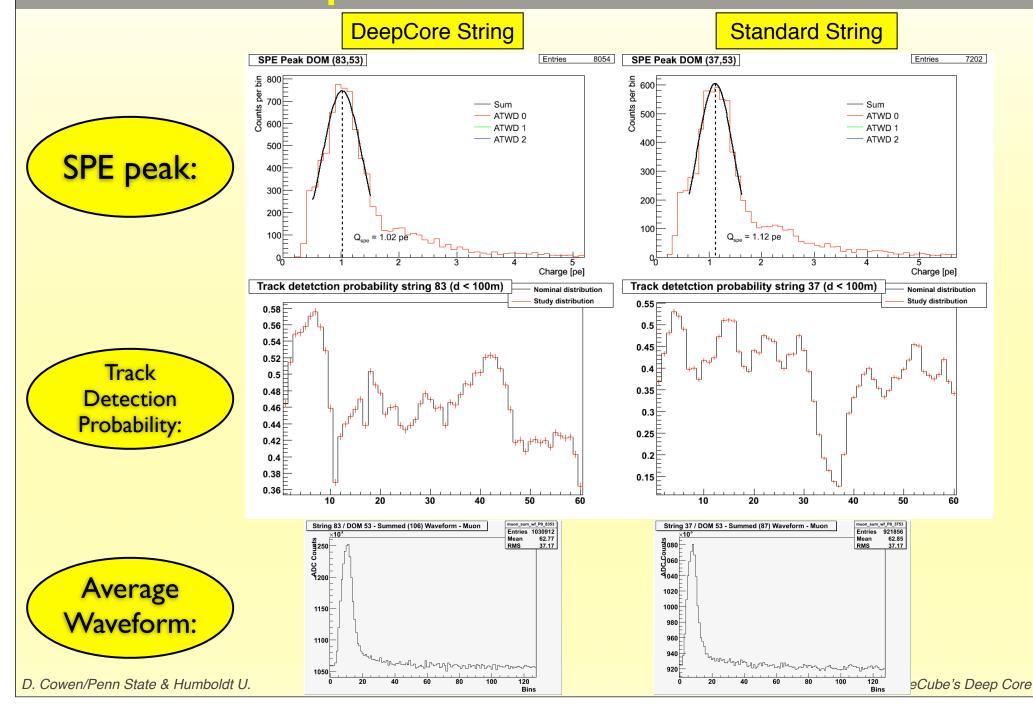
- First string already deployed; balance to be deployed next season
  - also, standard IceCube strings completing full veto will be deployed
- Performance of first string matches laboratory predictions for noise



# **DeepCore** Performance

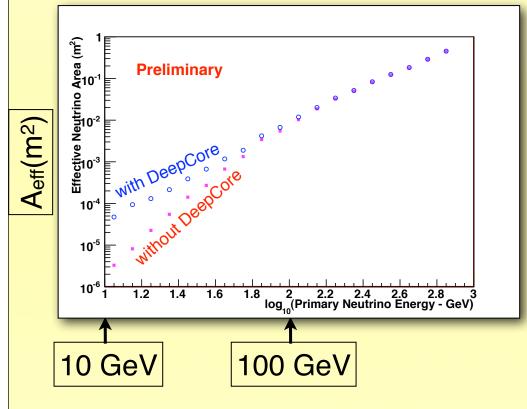


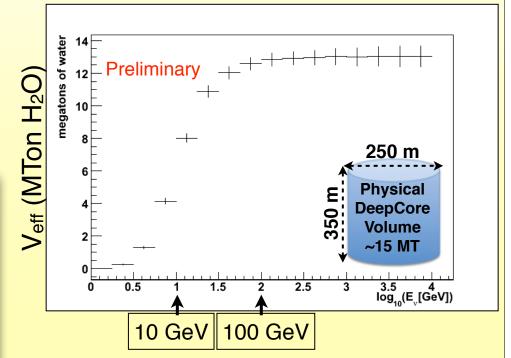
#### **DeepCore** Performance



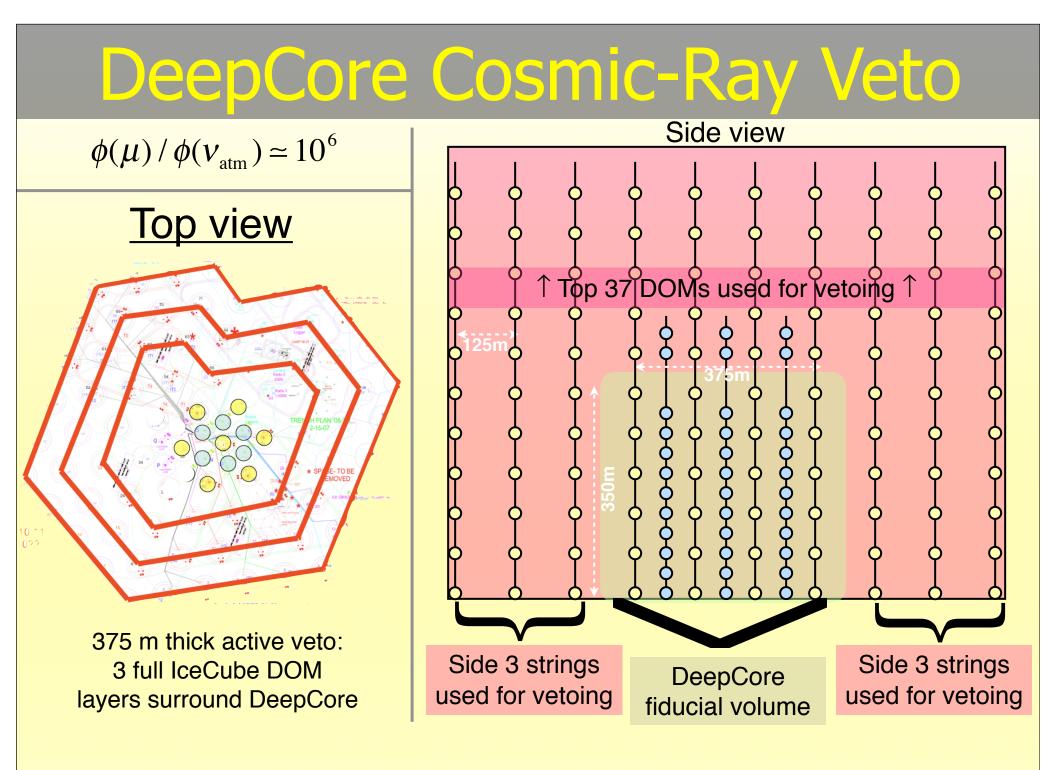
# DeepCore A<sub>eff</sub> and V<sub>eff</sub>

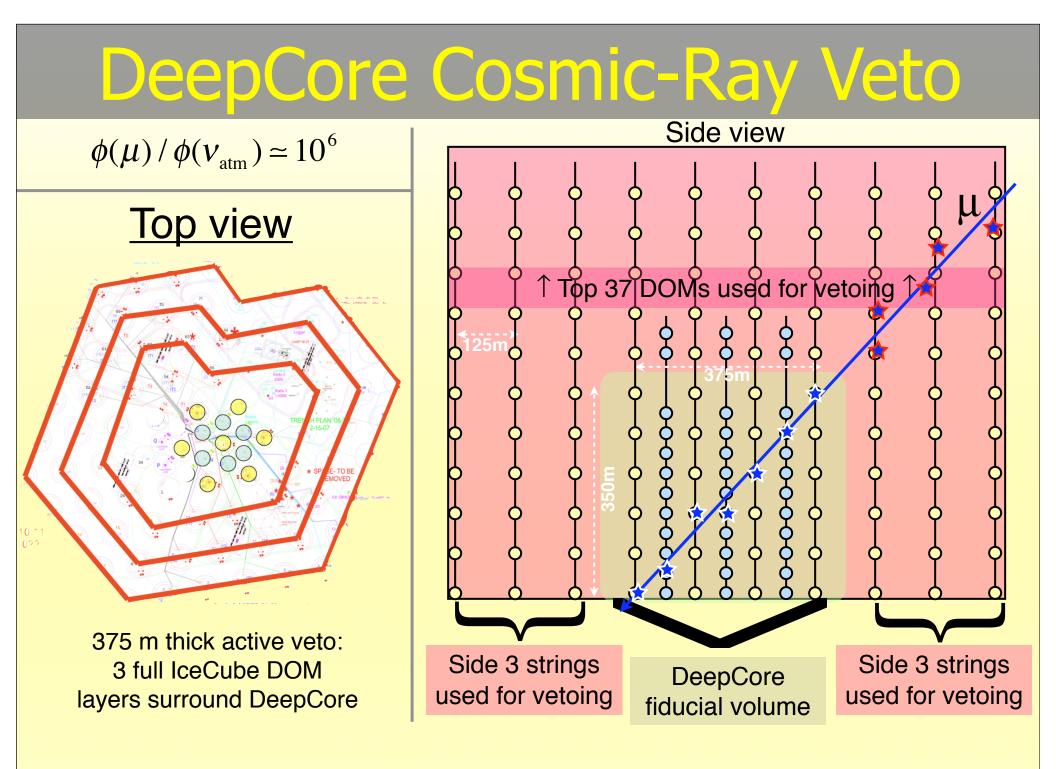
A<sub>eff</sub>: For downgoing muon neutrinos following E<sup>-2</sup> spectrum that trigger the detector ("SMT4", no reconstruction efficiencies included yet!)

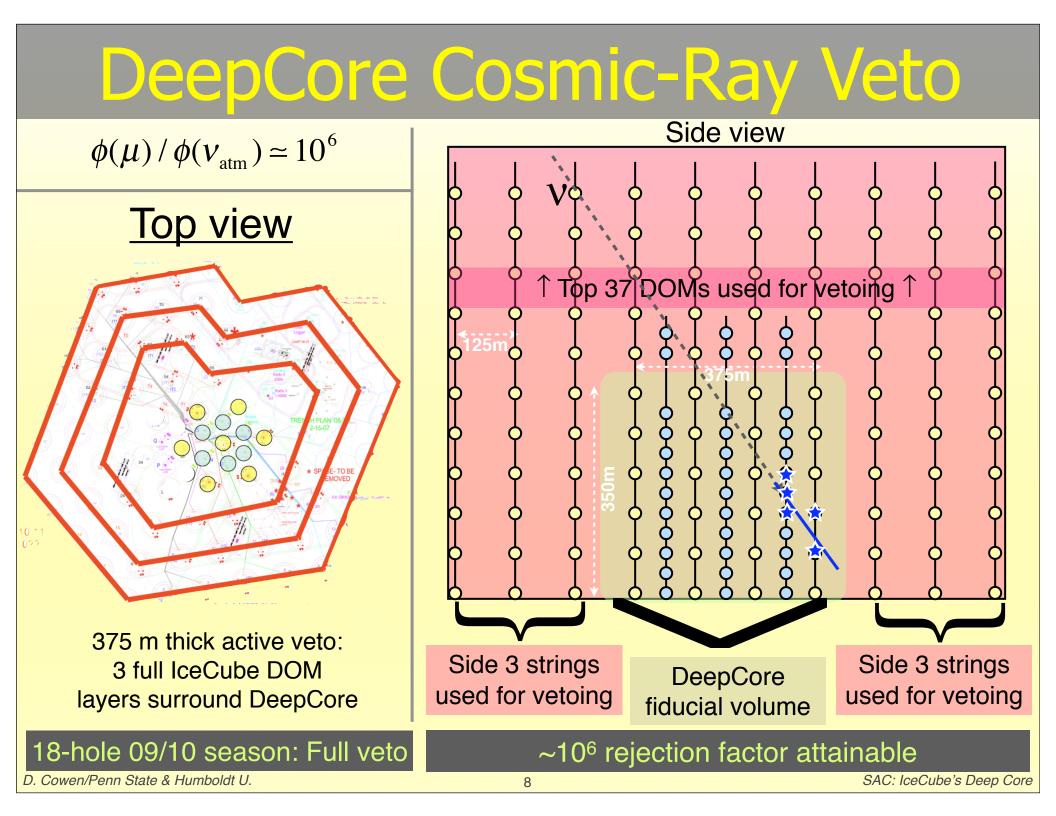




V<sub>eff</sub>: For contained downgoing muon neutrinos that interact in the fiducial volume and trigger the detector ("SMT4", no reconstruction efficiencies included yet!)

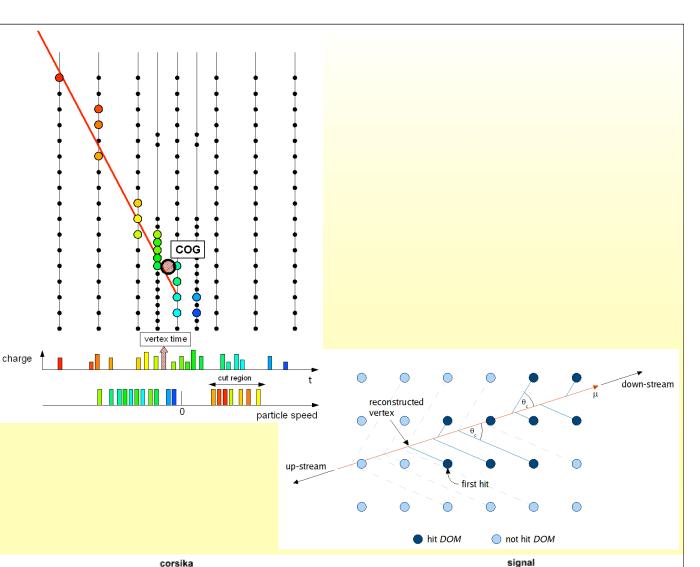






#### Veto Details

- Depth
- Veto events with hits outside fiducial volume consistent with a cosmic-ray muon
- Evaluate likelihood of event to have been produced by a neutrino vs. a muon
- Require event vertex to be in DeepCore fiducial volume



100 200

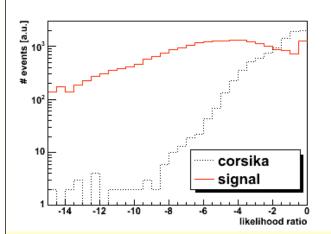
300 400 500

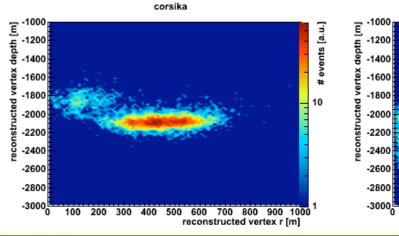
# events [a.u.]

10

600 700 800 900 1000

reconstructed vertex r [m]





# **DeepCore** Physics

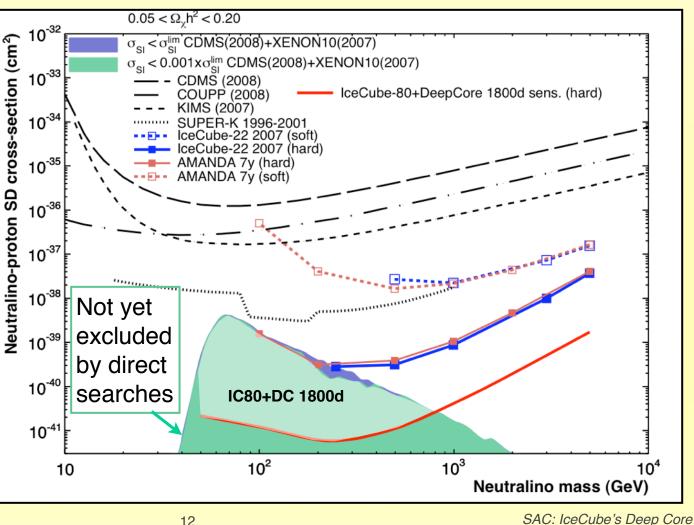
- Main topics:
  - WIMPs
  - Neutrino Oscillations
  - Southern Sky Sources
  - Exotica
- Studies performed with full detector simulation
  - WIMPs: basic reconstructions used
  - Oscillations: just at trigger level

# WIMPs and IC22

- A WIMP analysis was performed with IC22
- Achieved background rejection of 10<sup>6</sup>, signal efficiency of ~20%
  - 3° angular resolution
  - only used data with sun below horizon
- Observed flux with 104 days live time consistent with background expectation
  - background estimated from off-source data
- Set limit on  $\sigma_{SD}$  by assuming  $R_{annih} = R_{capture}$ , local  $\rho_{WIMP} = 0.3 GeV/cm^3$  and Maxwellian  $v_{WIMP}$

# WIMPs and DeepCore

- IC22 result improves existing limits on  $\sigma_{SD}$  (ArXiv: 0902.2460v1, accepted by PRL)
- Main systematic uncertainties (~20%):
  - photon propagation in the ice
  - absolute DOM efficiency
- DeepCore (& IC80) will extend sensitivity into region not yet excluded by direct searches based on σsi



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# **DeepCore and WIMPs**

- Significant improvement in expected solar WIMP sensitivity relative to last SAC meeting:
  - final geometry & high QE PMTs included
  - 10x more 1 $\mu$  and atm  $\nu$  background
  - coincident  $2\mu$  and  $3\mu$  background used
  - updated ice model
  - better trigger and updated reconstruction
  - improved analysis method (multivariate) and additional cuts
- N.B.: Our studies focused thus far on solar WIMPs
  - later will also study DeepCore response to Earth WIMPs and possibly WIMPs from galactic cente

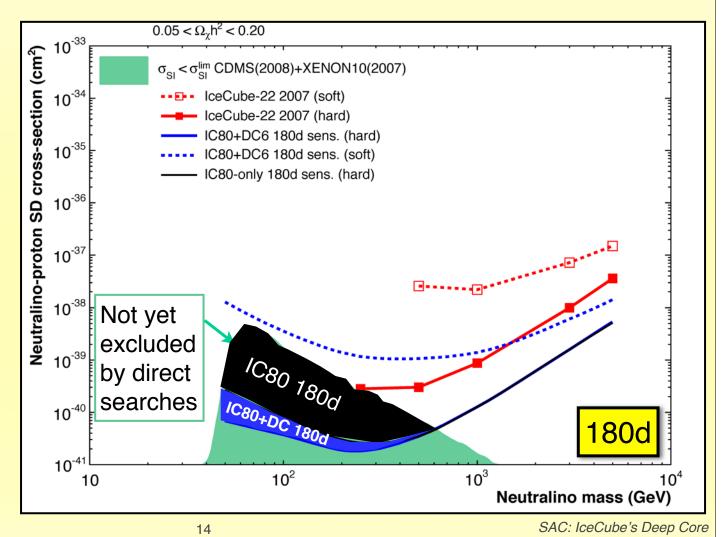
# WIMPs and DeepCore

 Same expected sensitivity as before, only in 1/10 the time

Very

 conservative
 DeepCore
 contribution at
 low energies

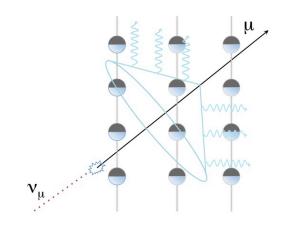
 minimizing reconstruction systematics etc. will be key to exploiting DeepCore's potential



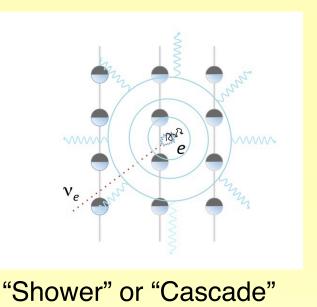
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## DeepCore & v Oscillations

- Preliminary studies performed using full detector simulation
  - assume high level of background suppression provided by veto
  - done only at trigger level
    - signal reconstruction algorithms to identify low energy neutrino flavor and energy are under development
- Three possible measurements
  - $v_{\mu}$  disappearance
  - $v_{\tau}$  appearance
  - neutrino hierarchy



"Track"

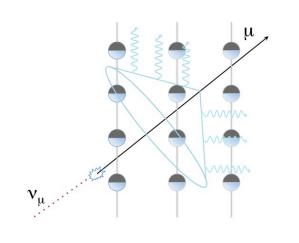


SAC: IceCube's Deep Core

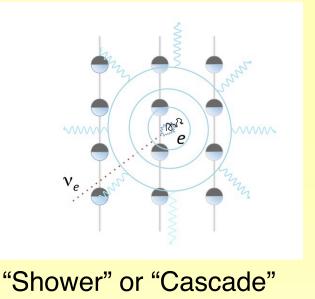
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  - $v_{\mu}$  disappearance
  - $v_{\tau}$  appearance
  - neutrino hierarchy

[Feasible.] [Maybe.] [<u>Very</u> hard.]

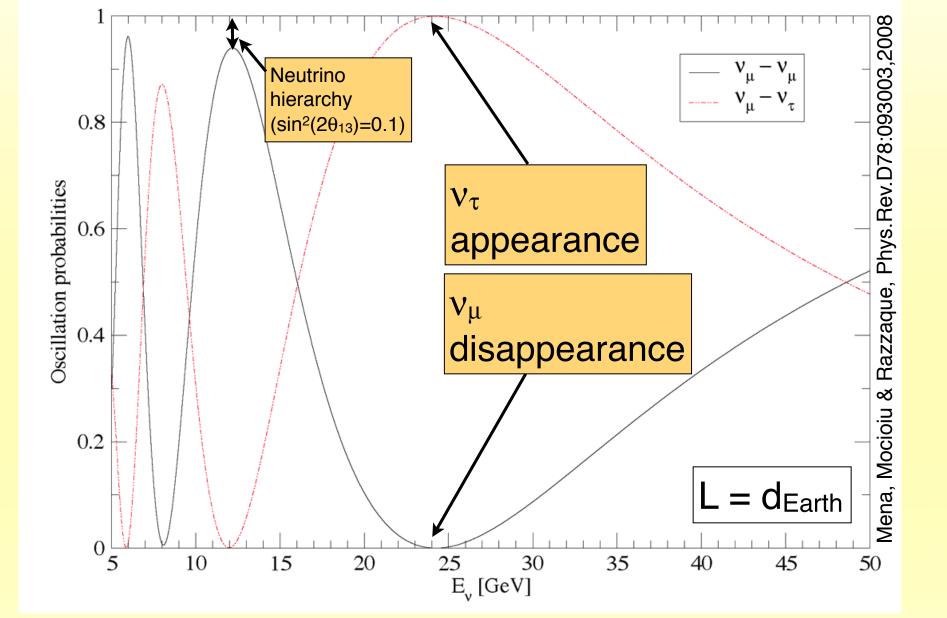


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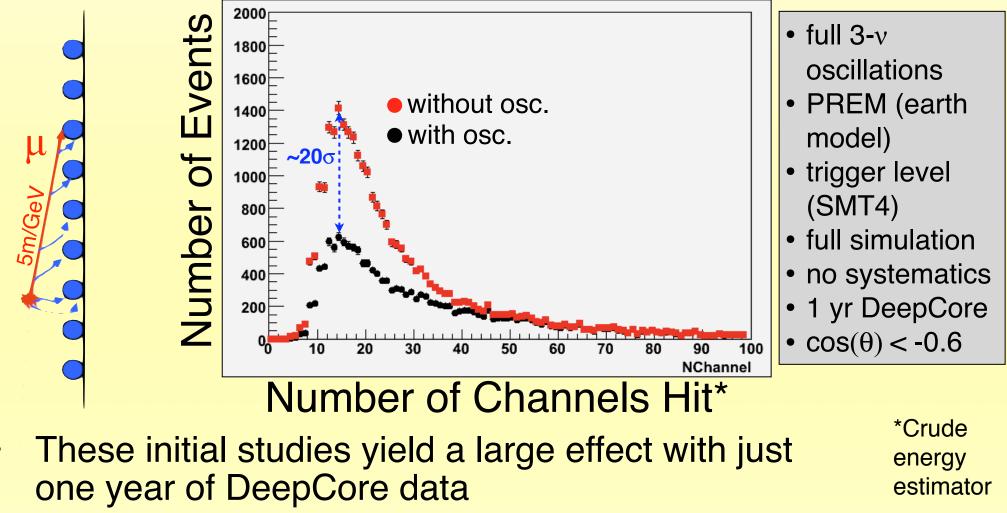
## DeepCore & v Oscillations



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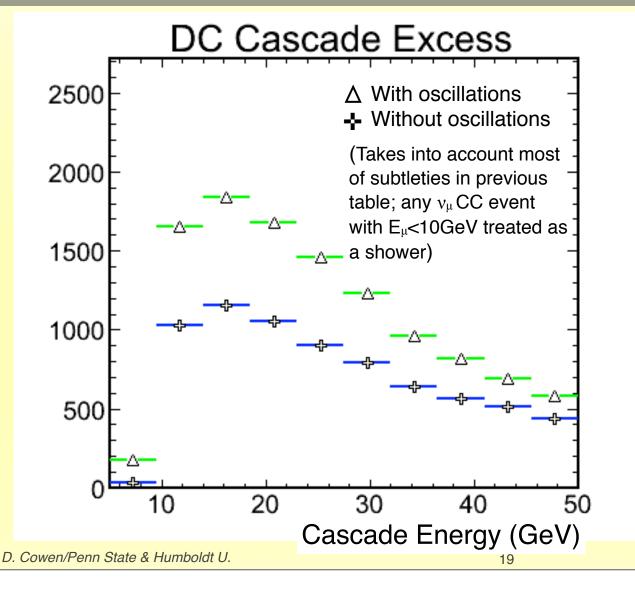
#### DeepCore & $v_{\mu}$ Disappearance

• Use ~vertically upgoing  $v_{\mu}$ -induced muons



SAC: IceCube's Deep Core

### v<sub>τ</sub> Appearance: Ultra-Preliminary Study

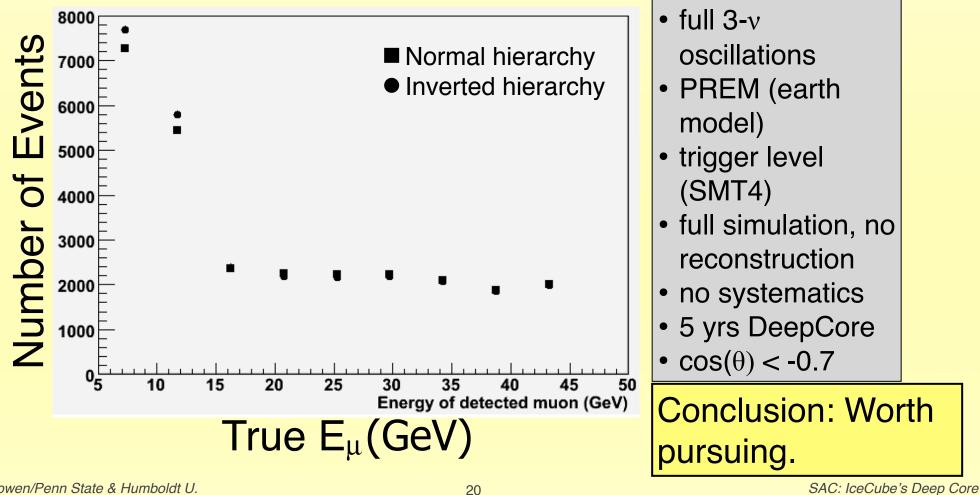


- full 3-v oscillations
- PREM (earth model)
- trigger level (SMT4)
- full simulation
- no systematics
- 1 yr DeepCore

Conclusion: Worth pursuing.

#### DeepCore & Hierarchy: Sensitivity

 Perform study with full IceCube simulation at trigger level (no reconstruction)

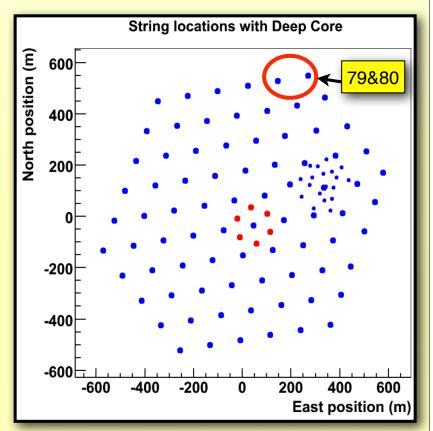


## Low Energy Reconstruction

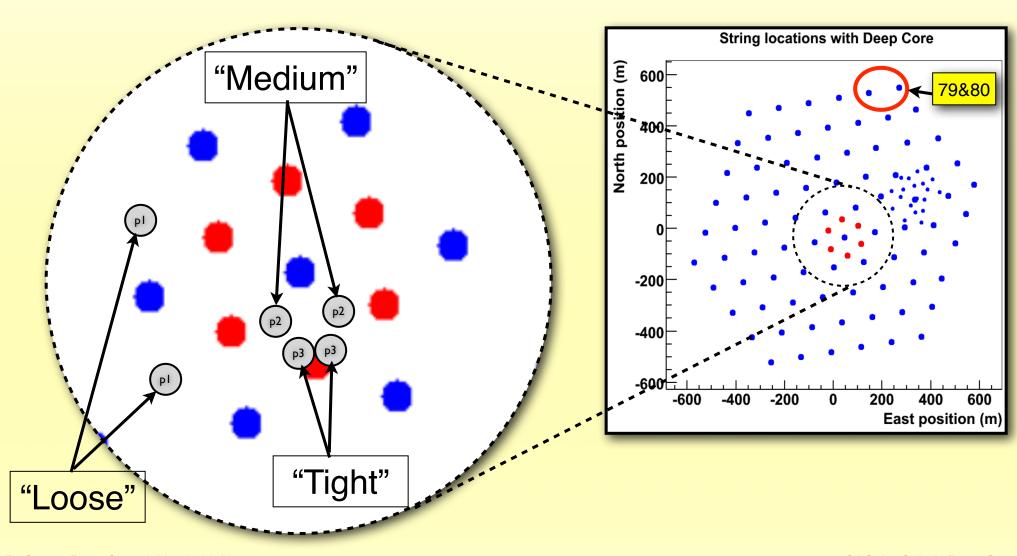
- Existing reconstruction algorithms for muon neutrinos assume "infinite" length
- At energies of interest to DeepCore, this is not a very good approximation
  - Must take into account both finite track length and hadronic shower at creation vertex
- Week-long mini-workshop planned for early June at Heidelberg to accelerate progress on this front
- Once have traction on reconstruction, will finish partially-completed DeepCore "detector paper"

## Strings 79 & 80

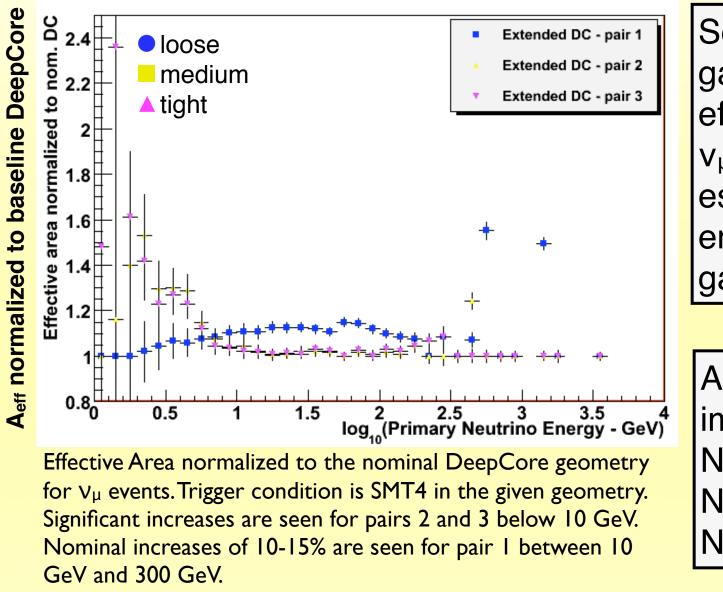
- We have performed a study of the advantages of adding two strings to the DeepCore array
  - studied 3 possible geometries
  - see marked improvement in:
    - basic reconstruction quantities
    - sensitivities
  - studies assumed strings would have the DeepCore DOM geometry
    - decision on breakouts before 1 July 2009
- Note: Any two strings could be used; 79&80 historically were first two available and started us thinking about them in DeepCore context



## Strings 79 & 80



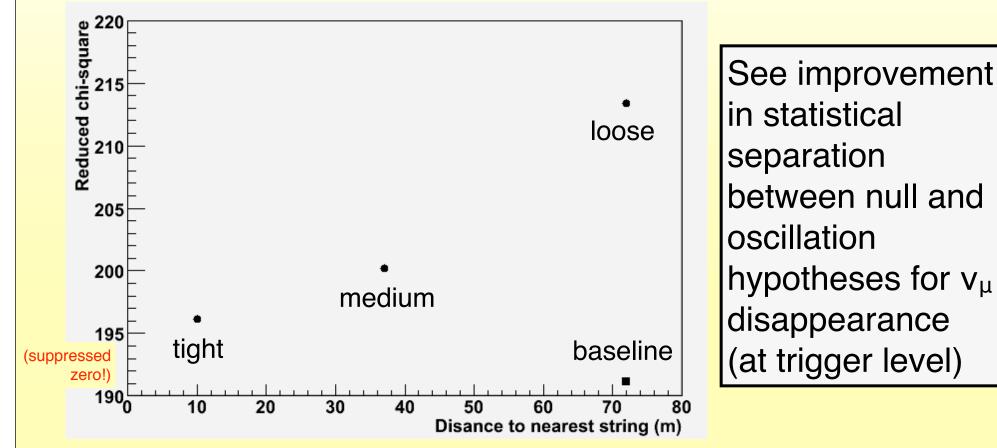
#### DeepCore + 79&80



See significant gain in the effective area for  $v_{\mu}$  events, especially at low energies. Similar gains for  $v_{e}$ .

Associated improvements: N<sub>str</sub>: up to 200% N<sub>ch</sub>: up to 30% N<sub>dir</sub>: up to 40%

### DeepCore + 79&80



Reduced chi-square calculated for the binned nchannel distributions of the expected (null hypothesis - no oscillations) and observed (numu disappearance) event rates. Trigger condition is SMT4 in the given detector geometry, normalized for 1 year of livetime. Circles- extended DeepCore geometries (pair 3 - 10m nearest string; pair 2 - 37m nearest string; pair 1 - 72m nearest string). Square - nominal DeepCore. Larger values of chi-square represent improved separation between the oscillation/no oscillation scenarios.

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# Disposition of 79&80

- We think we have a compelling case to use 79&80 to augment DeepCore
  - improves N<sub>str</sub>, N<sub>ch</sub>, and N<sub>dir</sub>, all markers for improved reconstruction capability
  - reconstruction is probably going to be the key challenge faced by low energy analyses
- However, strings 79 & 80 are also of interest to the High Energy Extension
- Have some indications that HEE works nearly as well without 79&80 as with (i.e., with 7 strings vs. 9)
  - concern: we may be near a "critical edge" and MC is not accurate enough to tell us

## Conclusions

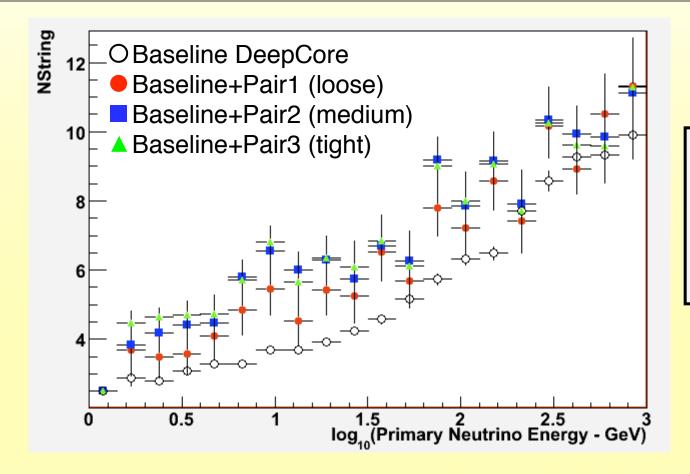
- Need to settle on location of strings 79&80 soon
  - SAC input requested
- Work on improving the S/N of the veto is ongoing
  - Cosmic-ray muon events tagged by IceTop can be used to estimate leakage systematics
- Efforts on new reconstruction algorithms tailored for low energy events are starting in earnest
- More exhaustive studies of the first DeepCore string 83 have begun:
  - need to fully characterize new PMTs (overall efficiency, charge response, linearity, noise, stability...)
  - need to understand photon propagation in deep ice better (not unique to DeepCore/Low Energy)

## The End

## Backup Slides

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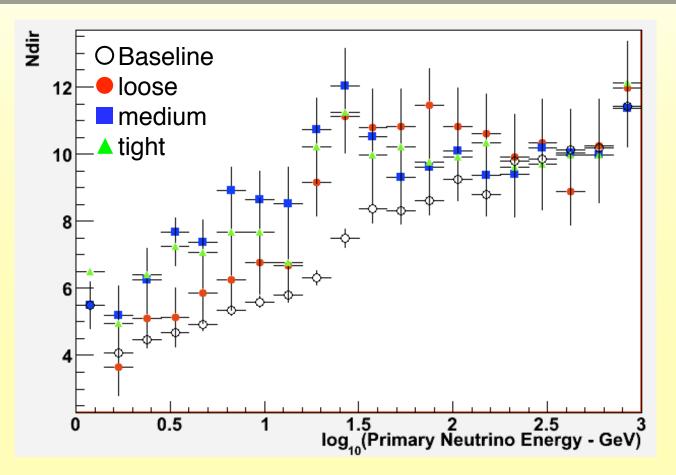
#### DeepCore + 79&80



See significant gain in the number of strings per event

Profile distribution for NString vs. Primary Neutrino Energy of NuMu events. Trigger condition is SMT4 in the given geometry. Pair I - red circles; Pair 2 - blue squares; Pair 3 - Green triangles; Nominal DeepCore - black open circles.

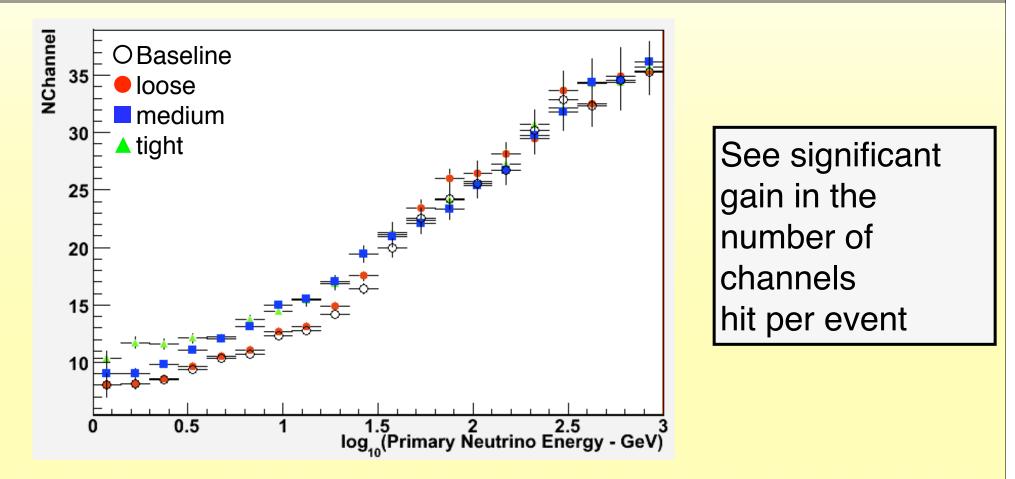
#### DeepCore + 79&80



See significant gain in the number of direct (unscattered) hits per event

Profile distribution for NDir vs. Primary Neutrino Energy of NuMu events. Trigger condition is SMT4 in the given geometry. Pair I - red circles; Pair 2 - blue squares; Pair 3 - Green triangles; Nominal DeepCore - black open circles. Used MC truth for track position and direction.

#### DeepCore + 79&80

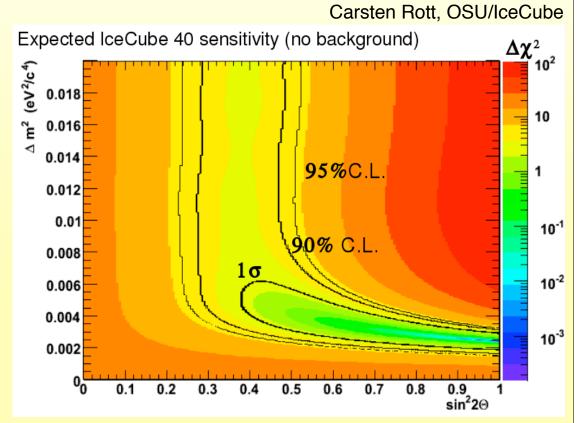


Profile distribution for NChannel vs. Primary Neutrino Energy of NuMu events. Trigger condition is SMT4 in the given geometry. Pair I - red circles; Pair 2 - blue squares; Pair 3 - Green triangles; Nominal DeepCore - black open circles. v<sub>µ</sub> Disappearance: IceCube Preliminary Study Results

- 1.45 days of MC background yielded zero events
  - good, but need more MC statistics! ...Being generated
- Used a 12.85-day subsample of IC22 data (requiring ≥ 8 hits, not ≥ 5...yet)
  - expected 1.8 signal events without oscillations; 1.4 with oscillations
  - saw 3: reasonable agreement so far...
  - E<sub>v</sub> > ~25 GeV
- For full 200+ day IC22 dataset
  - predict 35.3 without oscillations; 27.7 with

## Next Steps: IC40, IC59

- Use data from larger IC40 detector
  - 3x more "central" strings
  - minimum 5 instead of 8 hits
- Expect about a 10-fold improvement in statistics:
  - 353 events without oscillations, 277 with: a 4σ effect (optimistic: 0 background assumed)

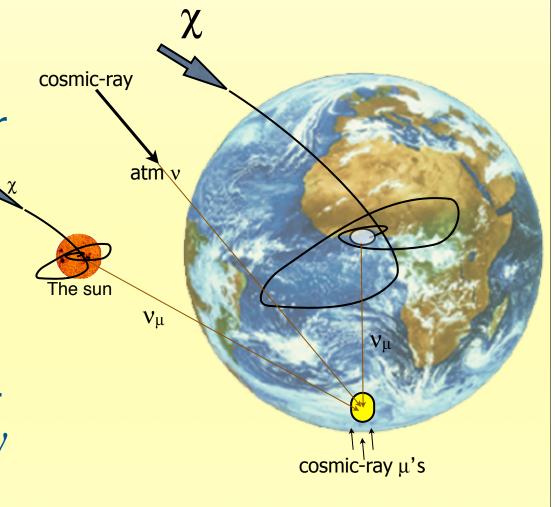


- Repeat with IC59
- Move on to DeepCore...

## Detecting a WIMP Signal

#### Look for:

- few signal events per year from the direction of the sun or earth
  - Soft:
    - $E_{\mu} \sim 0.01 M_{\chi} 0.06 M_{\chi}$
  - Hard:
    - $E_{\mu} \sim 0.03 M_{\chi} 0.3 M_{\chi}$
- Bkgd: ~5.10<sup>10</sup> cosmicray  $\mu$  and ~10<sup>5</sup> atm. v bkgd events per year



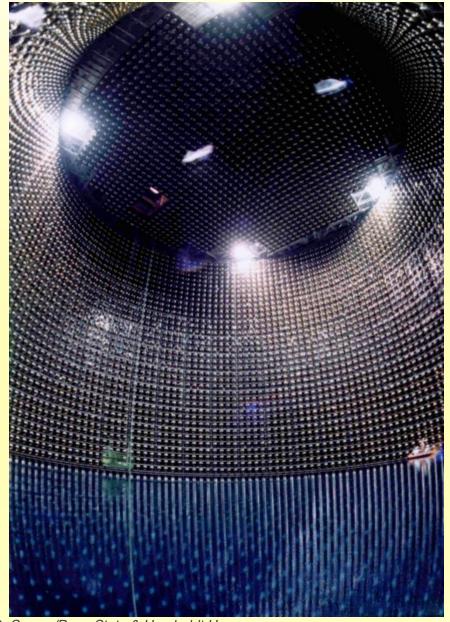
## WIMPs

WIMP annih • Earth's core • Solar core	<ul> <li>Indirect detection from WIMP– WIMP annihilation in the</li> <li>Earth's core</li> <li>Solar core</li> <li>Galactic center</li> </ul>	
$\begin{array}{ccc} q\bar{q} \\ \chi\chi \rightarrow & l\bar{l} & \rightarrow \cdots \rightarrow v_{\mu} \\ W^{\pm}, Z, H \end{array}$	<ul> <li>Indirect vs. direct searches:</li> <li>longer cosmological integration time</li> <li>sensitive to lower ν̄<sub>WIMP</sub></li> <li>sensitive to σ<sub>SD</sub> (solar)</li> </ul>	
$\chi\chi \rightarrow \overline{b}b(\tau^{+}\tau^{-}) \rightarrow \nu \qquad \text{soft } E_{\nu} \text{ spectrum}$ $\chi\chi \rightarrow W^{+}W^{-} \rightarrow \nu \qquad \text{hard } E_{\nu} \text{ spectrum}$		

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 $\chi$ 

#### v<sub>µ</sub> Disappearance: SuperK



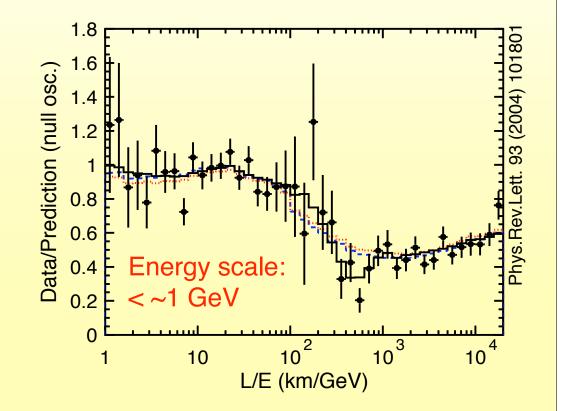
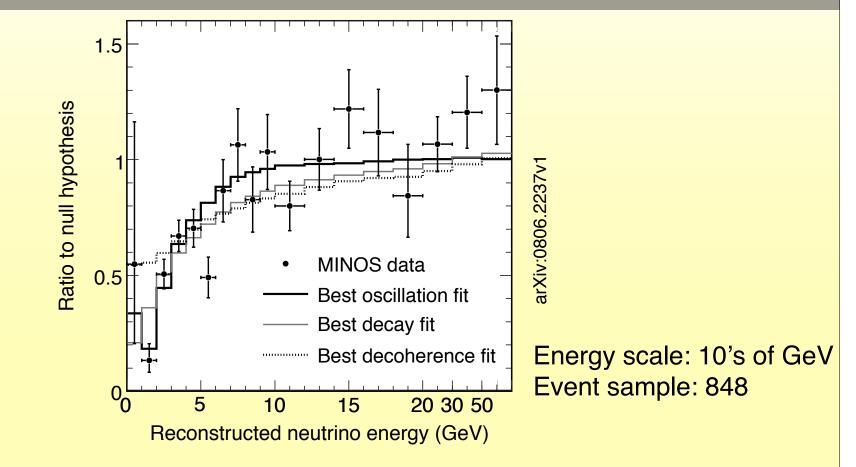


FIG. 4: Ratio of the data to the MC events without neutrino oscillation (points) as a function of the reconstructed L/E together with the best-fit expectation for 2-flavor  $\nu_{\mu} \leftrightarrow \nu_{\tau}$  oscillations (solid line). The error bars are statistical only. Also shown are the best-fit expectation for neutrino decay (dashed line) and neutrino decoherence (dotted line).

 $\mathsf{P}(v_{\alpha} \rightarrow v_{\beta}) = \sin^2 2\theta \sin^2 [1.27 \Delta m^2 (L/E)]$ 

#### v<sub>µ</sub> Disappearance: MINOS



# Measured: Δm<sup>2</sup> = (2.32±0.16 × 10<sup>-3</sup>) eV<sup>2</sup> and sin<sup>2</sup>2θ = 1.0

#### DeepCore & v Oscillations

- To be sensitive to ∆m<sup>2</sup>(atm) ~10<sup>-3</sup>, require L(km)/E(GeV)~10<sup>3</sup>
- At its design sensitivity of  $E_v \sim 1 \text{TeV}$ , IceCube needs  $L \sim 10^6 \text{ km}$ 
  - There are no TeV neutrino sources at that distance
- Atmospheric neutrinos, with L~10<sup>4</sup> km, exist in abundance and can be used by IceCube...

• ...but only if IceCube has sensitivity to  $E_v \sim 10 \text{ GeV}$ 

 [Note: Astrophysical neutrinos from distant sources give us sensitivity to oscillations, decay, ... but we need to find one or more sources first!]

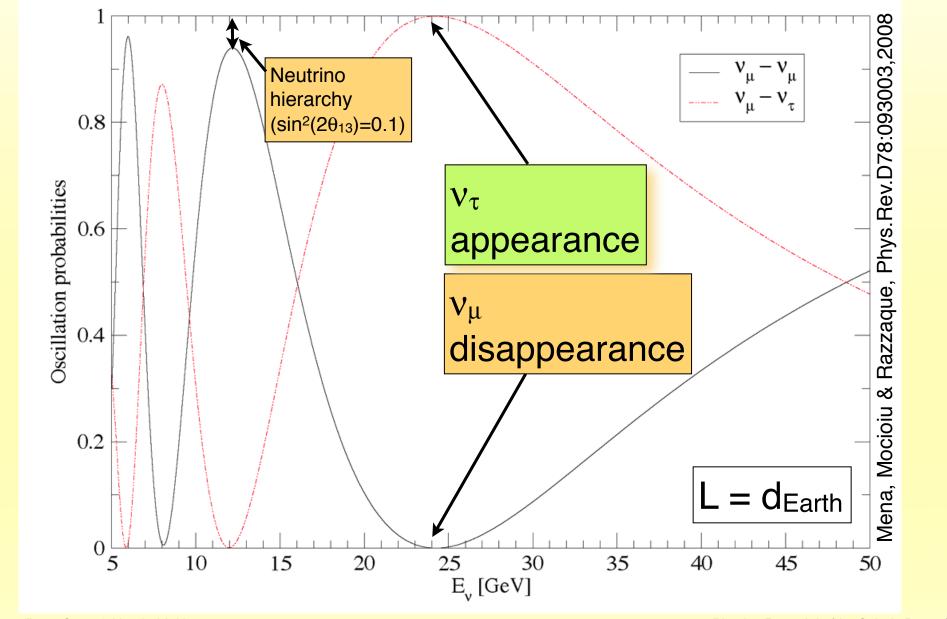
#### $v_{\tau}$ Appearance

- Indications are that ν<sub>µ</sub>→ν<sub>τ</sub>, but ν<sub>τ</sub> not yet seen directly
  - Difficult to measure:
    - $\tau$  has short lifetime, difficult to identify unambiguously
    - If try to compensate for short lifetime with larger E, then need very large L to get L/E in range
- CHORUS ('97) and NOMAD ('98) attempted to detect  $v_{\tau}$  from  $v_{\mu} \rightarrow v_{\tau}$  oscillations
  - placed limits
- SuperK "disfavors absence" of atmospheric  $v_{\mu} \rightarrow v_{\tau}$  at 2.4 $\sigma$
- OPERA (CNGS) has started running
  - expects ~2  $v_{\tau}$  per year

## $v_{\tau}$ Appearance: Technique

- Veto cosmic-ray background
- Reconstruct shower position and energy in DeepCore
- Use MC to correct for shower-like events from non-v<sub>τ</sub>-CC interactions
- Compare to null oscillation hypothesis, normalizing to higher energies

#### DeepCore & v Oscillations



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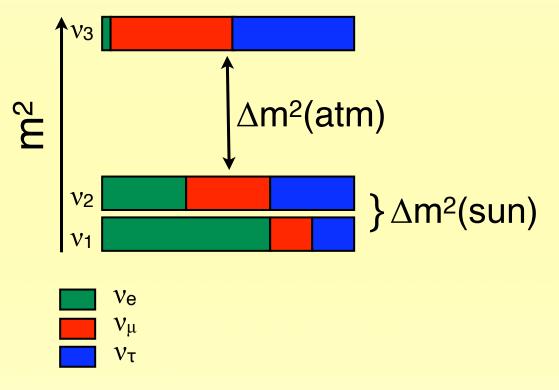
#### DeepCore & $v_{\tau}$ Appearance

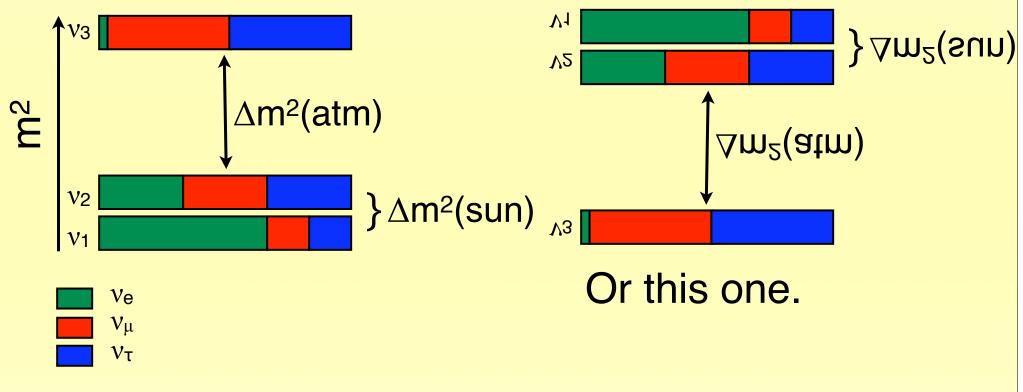
- Try to detect  $v_{\tau}$  signal from  $v_{\mu}$  oscillations
  - Look for enhancement in number of shower events around  $E_v \sim 25$  GeV (< $E_{vis} > \sim 10-15$  GeV)
  - Must distinguish shower-like from track-like events
- Showers are produced in many ways, and we must take into account all of them:

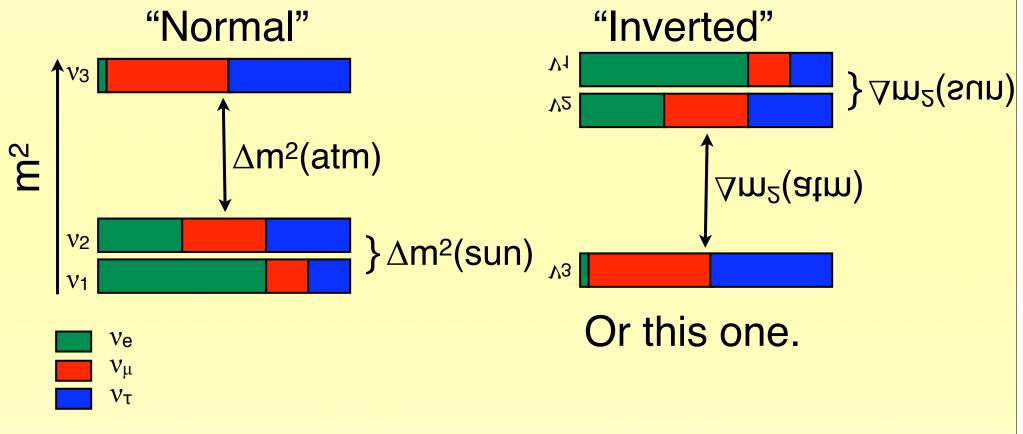
#### DeepCore & $v_{\tau}$ Appearance

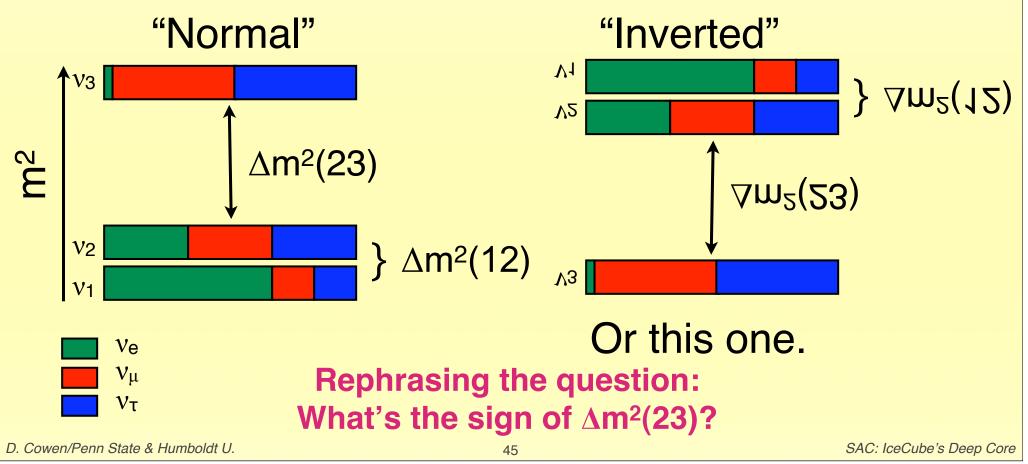
- Try to detect  $v_{\tau}$  signal from  $v_{\mu}$  oscillations
  - Look for enhancement in number of shower events around  $E_v \sim 25$  GeV (< $E_{vis} > \sim 10-15$  GeV)
  - Must distinguish shower-like from track-like events
- Showers are produced in many ways, and we must take into account all of them:

		CC	NC	Comment
	νe	Ew+Ee	Ez	$Φ(v_e)~0.1Φ(v_\mu);$ indep. of $v_\mu \rightarrow v_\tau$
	$ u_{\mu}$	veto(visible μ) OR <mark>Ew+E</mark> μ	Ez	CC & NC mimic signal; reduce significance
	$ u_{ au}$	E <sub>W</sub> +E <sub>e/μ/h</sub> -E <sub>(2/2/1 v's)</sub> OR veto(τ→μνν)	Ez	Will veto some $\tau$ signal that mimics CC $v_{\mu}$
m //	John Ctat	a 8 Humboldt II	10	SAC: losCuba's









- No experiment built so far can distinguish between "normal" and "inverted" hierarchies
  - need to be able to measure difference between  $\nu$  and  $\overline{\nu}$  interactions
    - rely on matter effects
- The NOvA experiment (under construction) may have sensitivity, provided Nature cooperates
  - Ditto for IceCube

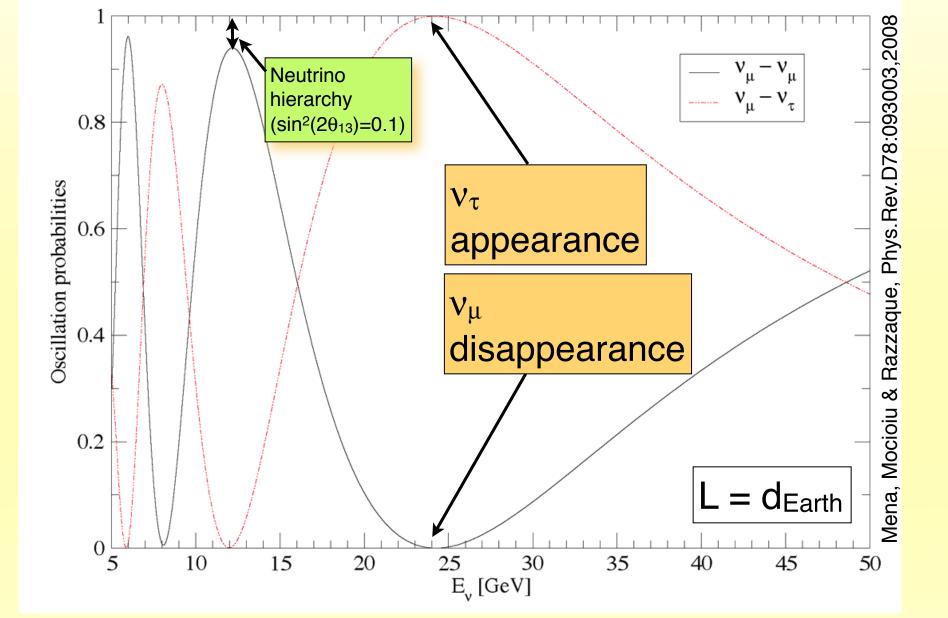
#### **DeepCore & Neutrino Hierarchy**

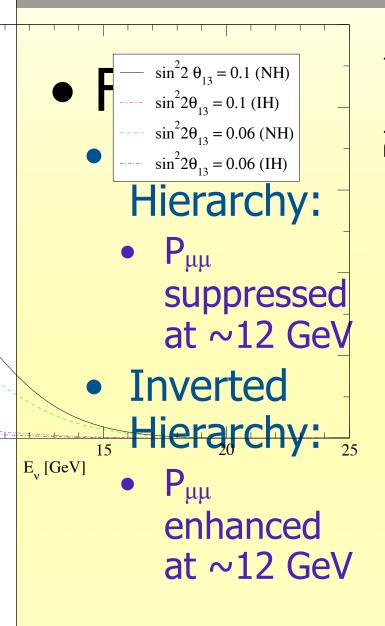
• Measurements use the fact that the effective  $\theta_{13}$  angle in matter is given by

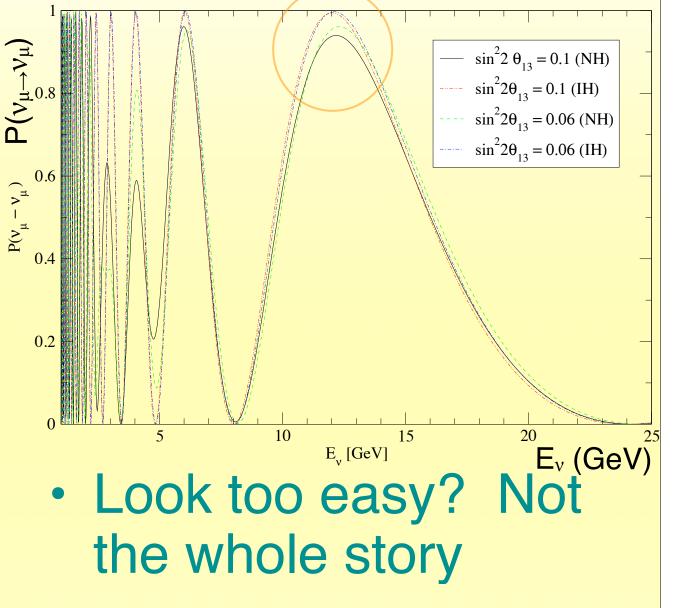
$$\sin^2 2\theta_{13}^{\rm m} = \frac{\sin^2 2\theta_{13}}{\sin^2 2\theta_{13} + \left(\cos 2\theta_{13} \mp \frac{\sqrt{2}G_F N_e}{\Delta_{31}}\right)^2}$$

- With "-" for v, "+" for  $\overline{v}$ ;  $\Delta_{31} \equiv (\Delta m_{31})^2 / 2E$
- At what  $E_v$  are we on resonance?
  - For  $L \sim d_E$ , with  $N_e$  for the earth:
    - $E_v$ (resonance) = ~ 10 GeV!
- So:
  - Look at low E, upgoing  $v_{\mu}$ -induced  $\mu$
  - Should see difference in disappearance for  $\nu_{\mu}$  vs.  $\overline{\nu}_{\mu}$

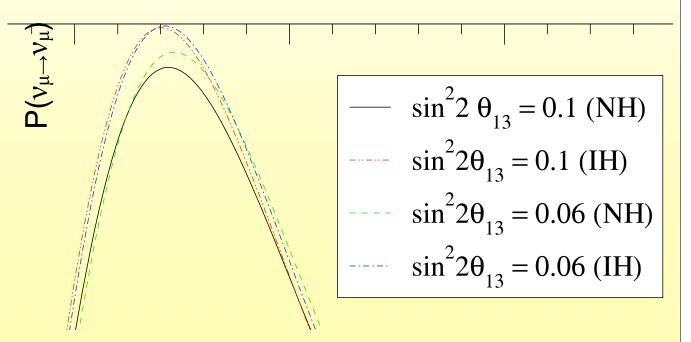
#### DeepCore & v Oscillations



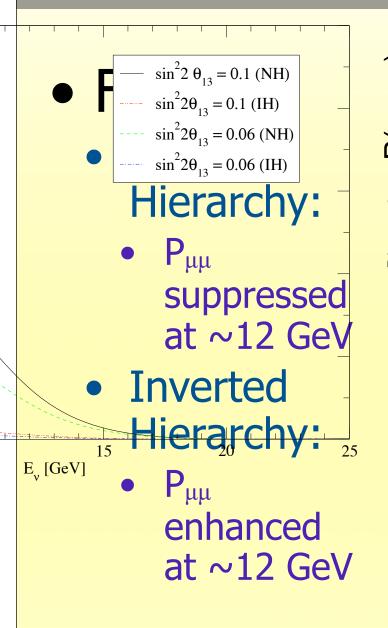


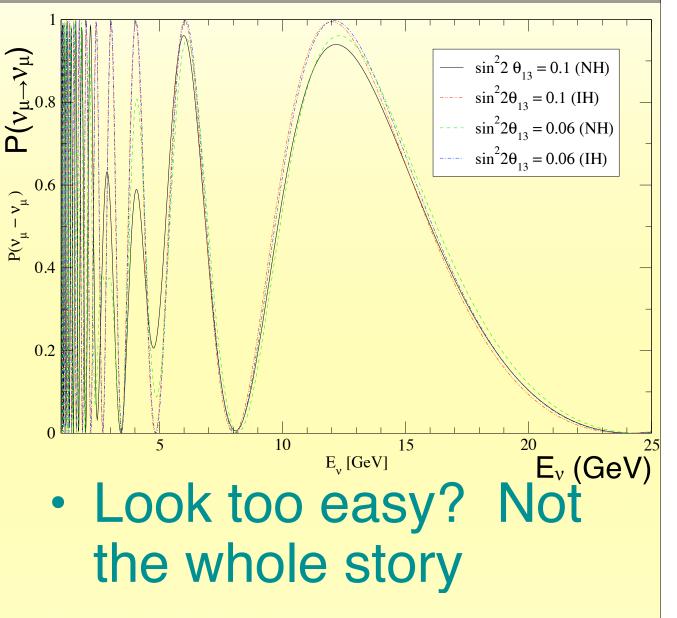


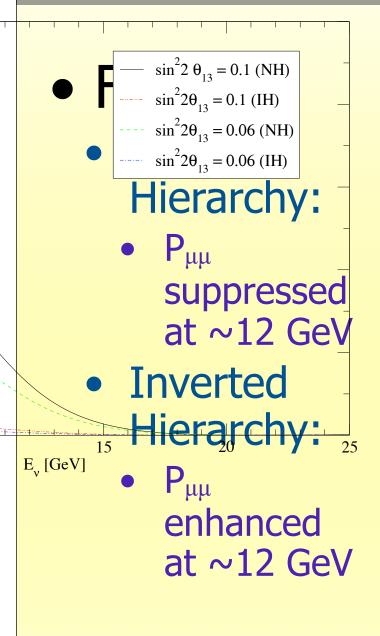
- For v's:
  - Normal Hierarchy:
    - P<sub>µµ</sub>
       suppressed
       at ~12 GeV
  - Inverted Hierarchy:
    - P<sub>µµ</sub>
       enhanced
       at ~12 GeV

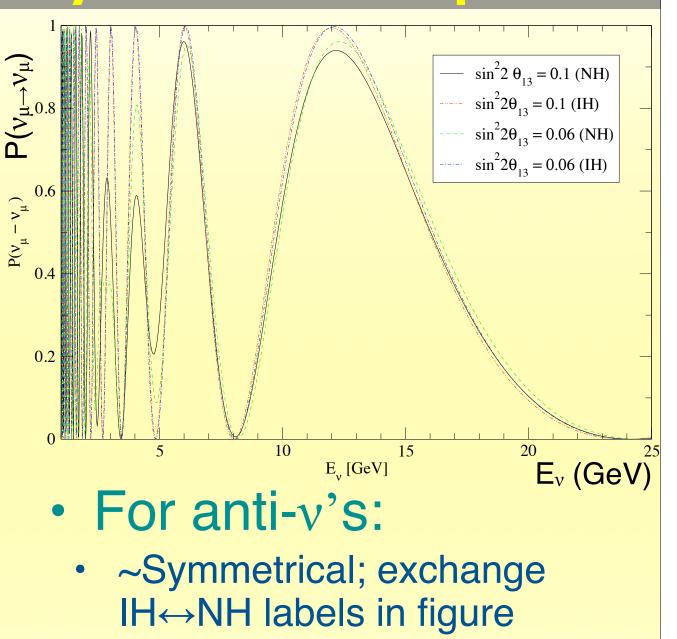


 Look too easy? Not the whole story





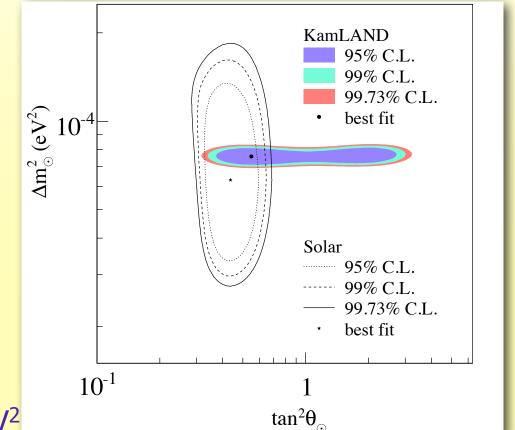




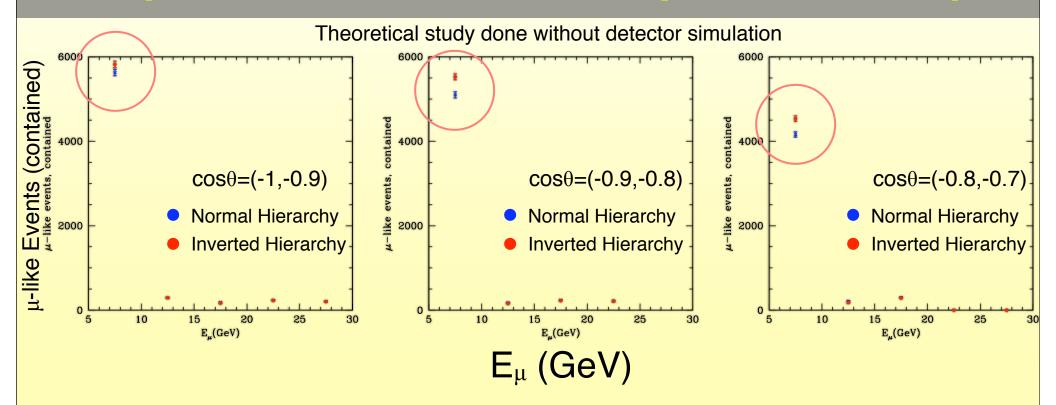
#### DeepCore & Hierarchy: Technique

- Can DeepCore tell the difference between  $v_{\mu}$  and  $\overline{v}_{\mu}$ ?
- It can't tell the difference between  $\mu^+$  and  $\mu^-...$
- But at E ~ 12 GeV:
  - $\sigma(v_x) \sim 2\sigma(\overline{v}_x)$ 
    - (effect diminished a bit by higher  $E_{\mu}$  of  $\overline{\nu}_{\mu}$  interactions)
- Thus we can get *statistical* discrimination between NH and IH

- (First, some preliminaries)
- Need to show the other main neutrino oscillation result, from solar and reactor neutrino experiments
  - These experiments measured  $v_e$  or  $\overline{v}_e$  disappearance
    - $\Delta m^2 = (7.6 \pm 0.2 \times 10^{-5}) \text{ eV}^2$
    - $tan^2\theta = 0.47 \pm 0.06$



#### DeepCore & Hierarchy: Technique

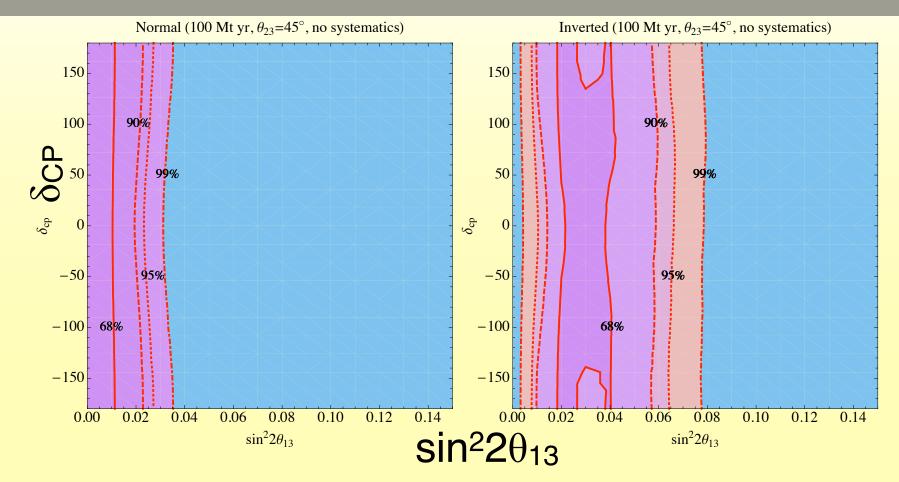


• 5 MTon detector, 10 yrs of running,  $E_{th} \sim 5 \text{ GeV}$ ,  $\theta_{13} = 0.1 (sin^2 2\theta_{13} = 0.04)^*$ 

- 5 GeV  $\mu$ ? L = 25 m, within 1.2 scattering lengths of 10-12 DOMs ( $\uparrow$ , centered)
- DeepCore physical volume is ~14 MTon
  - fiducial volume will be somewhere between 5 and 13 MTon

\*CHOOZ: sin<sup>2</sup>2θ<sub>13</sub> < 0.17 (90%CL) (Daya Bay/2011: sin<sup>2</sup>2θ<sub>13</sub> ~ 0.01)

#### **DeepCore & Hierarchy: Sensitivity**



#### Assumes 100 Mton-yrs

- Shows rejection level of wrong model
- Effect is easier to see if have NH due to difference in cross sections for  $\nu 's$  and  $\overline{\nu} 's$

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D. Cowen/Penn State & Normalia (1000/Mt yr,  $\theta_{23}$ =45°, 10% systematics)

Inverted (100 Mt yr,  $\theta_{23}$ =45°, 10% systematicSAC: IceCube's Deep Core